

(a) only

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9.17) Treat the linear antenna of problem 9.16, namely a thin linear antenna of length d excited so that the sinusoidal current makes one full ~~oscillation~~ wavelength of oscillation, by the multipole expansion method.

a) Calculate the multipole moments (electric dipole, magnetic dipole, and electric quadrupole) exactly and in the long-wavelength approximation.

Source current density is $\vec{J} = \hat{z} I \sin(kz) \delta(x) \delta(y) \Theta(\frac{d}{2} - |z|)$,

$$\text{for } k = \frac{2\pi}{\lambda} \text{ \& } \Theta(x) = \begin{cases} 1 & x > 0 \\ 0 & x < 0 \end{cases}$$

Exact expressions for multipole moments: (for the case $M=0$)

$$a_E(l, m) = \frac{k^2}{i\sqrt{l(l+1)}} \int d^3x Y_{lm}^* \left[c_P \frac{\partial}{\partial r} (r \hat{z} \cdot \vec{J}) + ik(\vec{r} \cdot \vec{J}) \hat{z} \right] \hat{z} \cdot \vec{J} \quad (9.167)$$

$$a_M(l, m) = \frac{k^2}{i\sqrt{l(l+1)}} \int d^3x Y_{lm}^* \nabla \cdot (\vec{r} \times \vec{J}) \hat{z} \quad (9.168)$$

Rewrite current density in spherical coordinates:

$$\vec{J} = \hat{z} I \sin(kz) \delta(x) \delta(y) \Theta(\frac{d}{2} - |z|)$$

$$= \hat{z} \frac{I}{2\pi r^2} \sin(kr) \left[\delta(\cos\theta - 1) + \delta(\cos\theta + 1) \right] \Theta(\frac{d}{2} - r)$$

From the continuity eqn's,

$$\rho = \frac{1}{i\omega} \nabla \cdot \vec{J} = -\frac{iI}{2\pi r^2 \omega} k \cos(kr) \left[\delta(\cos\theta - 1) + \delta(\cos\theta + 1) \right] \Theta(\frac{d}{2} - r)$$

(+ contribution from δ'' 's & endpoints which we omit)

$$\text{Use } k = \frac{\omega}{c}$$

9.17 a), cont'd

For later use,

$$\vec{r} \cdot \vec{J} = \frac{I}{2\pi r} \sin(kr) \left[\delta(\cos\theta - 1) + \delta(\cos\theta + 1) \right]$$

$$\vec{r} \times \vec{J} = 0$$

$$\Rightarrow \boxed{q_{lm}(l, m) = 0} \text{ for all } l, m$$

$$q_E(l, m) = \frac{-k^2 I}{\sqrt{l(l+1)}} \int d^3x$$

$$q_E(l, m) = \frac{-k^2 I}{\sqrt{l(l+1)}} \int d^3x \left[\frac{\cos kr}{2\pi r^2} \frac{\partial}{\partial r} (r j_l(kr)) - \frac{k}{2\pi r} \sin(kr) j_l(kr) \right]$$

$$\int \left[\delta(\cos\theta - 1) + \delta(\cos\theta + 1) \right] Y_{lm}^* r^2 dr d\Omega$$

$$= \frac{-k^2 I}{\sqrt{l(l+1)}} \delta_{m0} \left[Y_{l0}^*(0,0) + Y_{l0}^*(\pi,0) \right]$$

$$\int_0^{d/2} dr \left[\frac{\cos kr}{2\pi r^2} \frac{\partial}{\partial r} (r j_l(kr)) - \frac{k}{2\pi r} \sin(kr) j_l(kr) \right]$$

$$= \frac{\partial}{\partial r} (r \cos(kr) j_l(kr))$$

$$= \frac{-k^2 I}{\sqrt{l(l+1)}} \delta_{m0} \left[Y_{l0}^*(0,0) + Y_{l0}^*(\pi,0) \right] \left[r \cos(kr) j_l(kr) \right] \Big|_0^{d/2}$$

$$= \frac{-k^2 I}{\sqrt{l(l+1)}} \delta_{m0} \left[Y_{l0}^*(0,0) + Y_{l0}^*(\pi,0) \right] \frac{d}{2} \cos\left(\frac{d}{2}\right) j_l\left(\frac{d}{2}\right) \text{ using } k = \frac{2\pi}{d}$$

9.17 a), cont'd

$$\text{Use } Y_{\ell 0}(\theta, \phi) = \left(\frac{2\ell+1}{4\pi}\right)^{1/2} P_{\ell}(\cos\theta) \quad (9.57)$$

$$\& P_{\ell}(0) = 1, \quad P_{\ell}(-x) = (-1)^{\ell} P_{\ell}(x)$$

$$\Rightarrow a_E(\ell, m) = \left[-\left(\frac{2\pi}{d}\right)^2 \frac{d}{2} \int \left(\frac{2\ell+1}{\ell(\ell+1)}\right)^{1/2} \frac{1}{\sqrt{4\pi}} (1+(-)^{\ell}) (-)^m j_{\ell}(\pi) \delta_{m0} \right. \\ \left. = + \frac{2\pi}{d} \int \left[\frac{(2\ell+1)\pi}{\ell(\ell+1)}\right]^{1/2} \left(\frac{1}{2}\right) (1+(-)^{\ell}) j_{\ell}(\pi) \delta_{m0} \right] \quad (\text{exact answer})$$

Note that $a_E(\ell, m) = 0$ for ℓ odd,

so there is no electric dipole contribution.

Next, consider long-wavelength approximation.
Hence,

$$a_E(\ell, m) \approx \frac{ck^{\ell+2}}{i(2\ell+1)!!} \left(\frac{\ell+1}{\ell}\right)^{1/2} (a_{\ell m} + a'_{\ell m}) \quad (9.69)$$

where

$$\left. \begin{aligned} a_{\ell m} &= \int d^3x r^{\ell} \rho Y_{\ell m}^* \\ a'_{\ell m} &= \frac{-ih}{(\ell+1)c} \int d^3x r^{\ell} Y_{\ell m}^* \nabla \cdot (\mathbf{E} \times \mathbf{p}) \end{aligned} \right\} (9.170)$$

9.17 a), cont'd

Here, $Q'_{em} = 0$ since $\bar{M} = 0$

$$\begin{aligned}
 Q_{em} &= \int d^3x Y_{em}^* r^l \frac{(-i)I}{2\pi r^2 c} \omega(kr) \left[\delta(\cos\theta - 1) + \delta(\cos\theta + 1) \right] \theta\left(\frac{a}{2} - r\right) \\
 &= \frac{-iI}{c} S_{m0} \left[Y_{e0}^*(0,0) + Y_{e0}^*(\pi,0) \right] \int_0^{dr} dr (\cos kr) r^l \\
 &= \frac{-iI}{c} S_{m0} \left(\frac{2l+1}{4\pi} \right)^{1/2} \left(P_l(1) + P_l(-1) \right) \int_0^\pi dx (\cos x) x^l k^{-l-1} \\
 &= \frac{-iI}{c} S_{m0} \frac{1}{k^{l+1}} \left(\frac{2l+1}{4\pi} \right)^{1/2} (1 + (-1)^l) \int_0^\pi dx x^l (\cos x)
 \end{aligned}$$

$$\text{For } l=0, \int_0^\pi dx x^l \cos x = +\sin x \Big|_0^\pi = 0$$

so the first nonvanishing contribution is at $l=2$, consistent with the exact result.

For $l=2$,

$$\begin{aligned}
 \int_0^\pi dx \frac{x^2}{u} \frac{d}{dv} \cos x &= +x^2 \sin x \Big|_0^\pi - \int_0^\pi dx 2x (+\sin x) \\
 &= 2 \int_0^\pi \frac{x \sin x}{u} dx \\
 &= -2 \left[x(-\cos x) \Big|_0^\pi - \int_0^\pi dx (-\cos x) \right] \\
 &= -2 \left[\pi + \sin x \Big|_0^\pi \right] \\
 &= -2\pi
 \end{aligned}$$

$$\Rightarrow Q_{2m} = -i \frac{I}{c} S_{m0} \frac{1}{k^3} \left(\frac{5}{4\pi} \right)^{1/2} (2) (-2\pi)$$

9.17 a), cont'd

In any event, in the long-wavelength approximation,

$$\begin{aligned}
 a_E(l, m) &\approx \frac{c h^{l+2}}{i (2l+1)!!} \left(\frac{l+1}{l}\right)^{1/2} (-i) \frac{I}{c h^{l+1}} \delta_{m0} \left(\frac{2l+1}{4\pi}\right)^{1/2} (1+(-)^l) \int_0^\pi dx x^l \cos x \\
 &= -\frac{kI}{(2l+1)!!} \left[\frac{(l+1)(2l+1)}{4\pi l}\right]^{1/2} (1+(-)^l) \int_0^\pi dx x^l \cos x \\
 &= -\frac{2\pi}{d} \frac{I}{(2l+1)!!} \left[\frac{(l+1)(2l+1)}{4\pi l}\right]^{1/2} (1+(-)^l) \int_0^\pi dx x^l \cos x
 \end{aligned}$$

Similarly, in the long-wavelength approximation,

$$a_M(l, m) \approx \frac{i k^{l+2}}{(2l+1)!!} \left(\frac{l+1}{l}\right)^{1/2} (M_{em} + M'_{em}) \quad (9.171)$$

where

$$M_{em} = -\frac{1}{l+1} \int d^3x r^l y_{lm}^* \nabla \cdot (\mathbf{r} \times \mathbf{J}) \quad (9.172)$$

$$M'_{em} = -\int d^3x r^l y_{lm}^* \nabla \cdot \bar{\mathbf{m}}$$

but since both $\bar{\mathbf{m}} = 0$ & $\mathbf{r} \times \mathbf{J} = 0$,

we see

$$\boxed{a_M(l, m) = 0} \quad \text{in the long-wavelength approximation,}$$

consistent with the exact answer.

(a) only

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9.22 A spherical hole of radius a in a conducting medium can serve as an electromagnetic resonant cavity.

a) Assuming infinite conductivity, determine the transcendental equations for the characteristic frequencies ω_{em} of the cavity for TE and TM modes.

TE: $\vec{r} \cdot \vec{E} = 0$ (= magnetic multipole)

$$\left. \begin{aligned} \vec{E}_{em}^{(M)} &= z_0 j_2(kr) \vec{X}_{em} \\ \vec{H}_{em}^{(M)} &= -\frac{i}{kz_0} \nabla \times \vec{E}_{em}^{(M)} \\ &= -\frac{i}{k} \nabla \times (j_2(kr) \vec{X}_{em}) \end{aligned} \right\} (9.116) \quad \begin{array}{l} \text{w/ } j_2 \text{ taken to be } j_2, \\ \text{to ensure} \\ \text{well-behaved at origin} \end{array}$$

Perfect conductor $\Rightarrow H_{\perp} = 0, E_{\parallel} = 0$ at $r = a$.

$$\Rightarrow \vec{r} \cdot \vec{H}|_{r=a} = 0 = \vec{r} \times \vec{E}|_{r=a}$$

Both of these $\propto j_2(kr)$, hence $j_2(ka) = 0$

$$\Rightarrow k = \frac{x_{en}}{a}, \text{ where } x_{en} \text{ is the } n^{\text{th}} \text{ root of } j_2$$

$$\Rightarrow \boxed{\text{TE frequencies are } \omega = \frac{x_{en} c}{a}} \quad \text{where } j_2(x_{en}) = 0$$

(note each l frequency has degeneracy $2l+1$,
from possible m values,

9.22 a), cont'd

TM: $\vec{r} \cdot \vec{H} = 0$ (= electric multipoles)

$$\left. \begin{aligned} \vec{H}_{em}^{(E)} &= \hat{j}_e(kr) \vec{X}_{em} \\ \vec{E}_{em}^{(E)} &= \frac{i\epsilon_0}{k} \nabla \times \vec{H}_{em}^{(E)} \end{aligned} \right\} (9.118)$$

Perfect conductor $\Rightarrow H_{\perp} = 0, E_{\parallel} = 0$ at $r=a$

$H_{\perp} = 0$ automatic since $\vec{r} \cdot \vec{X}_{em} = 0$

$$E_{\parallel} = 0 \Rightarrow \vec{r} \times \left(\nabla \times (\hat{j}_e(kr) \vec{X}_{em}) \right) = 0 \text{ at } r=a$$

Write

$$\vec{r} \times (\nabla \times \vec{A}) = \nabla(\vec{r} \cdot \vec{A}) - (\vec{r} \cdot \nabla) \vec{A} - \underbrace{(\vec{A} \cdot \nabla) \vec{r}}_{\vec{A}} - \underbrace{\vec{A} \times (\nabla \times \vec{r})}_{=0} \quad (\text{point } \vec{A} \text{ over } \vec{r})$$

$$= \nabla(\vec{r} \cdot \vec{A}) - (\vec{r} \cdot \nabla) \vec{A} - \vec{A}$$

$$= \nabla(\vec{r} \cdot \vec{A}) - (\vec{r} \cdot \nabla) \vec{A} - \vec{A}$$

$$= \nabla(\vec{r} \cdot \vec{A}) - \left(1 + r \frac{\partial}{\partial r}\right) \vec{A}$$

$$= \nabla(\vec{r} \cdot \vec{A}) - \frac{\partial}{\partial r} (r \vec{A})$$

Here, $\vec{A} = \hat{j}_e(kr) \vec{X}_{em}$

Since $\vec{r} \cdot \vec{X}_{em} = 0, \vec{r} \cdot \vec{A} = 0$

$$\Rightarrow \vec{r} \times \left(\nabla \times (\hat{j}_e(kr) \vec{X}_{em}) \right) = -\frac{\partial}{\partial r} (r \hat{j}_e(kr)) \vec{X}_{em}$$

9.22 a), cont'd

Thus, $E_{11} = 0 \Rightarrow \boxed{\omega = \frac{\gamma_{en} c}{a}}$ are TM frequencies

where $\frac{d}{dx} [x j_2(x)] \Big|_{x=\gamma_{en}} = 0$ defines γ_{en}